

Design and Use of the KRC Thermal Model



Hugh H. Kieffer
Celestial Reasonings
hkieffer@charter.net
775-782-0767

Observe temperatures, but commonly want physical properties. Connected by a model. $I = \sqrt{k\rho C}$ = Thermal Inertia

- $K=k$ = conductivity

- $R=\rho$ = ρ =density

- $C=$ specific heat

- Basic temperature response to a periodic radiation boundary condition

Required reading: JGR-Planets 2013

Thermal model for analysis of Mars infrared mapping

Purpose: Transfer knowledge of KRC

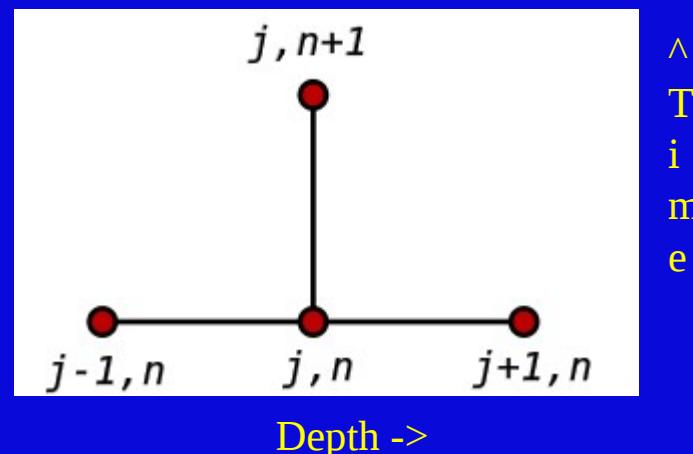
- A. Users of the maintained version on website: krc.mars.asu.edu
- B. Folks wanting their own copy: at ASU or elsewhere
- C. People wishing to modify KRC at their site
- D. Curation staff:
 - Cognizant programmer
 - Mentor
 - Website master
- Documentation is not always reliable. Search the code

Outline

- A brief history of a long time
- Physics included and omitted
- Input format: normal and OnePoint
- Examples
- Architecture
 - Run, case, season, latitude, day, time-step, layer
 - Call Diagrams: Flow between routines
- Use of Commons
- Output formats: Print and binary
 - Sequential Access: 3 each
 - .bin5 packed: 5 each
- PORB: Objects other than Mars
- Liens: Things yet to be done
- [[Typical Use: Model Sets; detailed studies

History

- Begun in 1968 to support Mariner 6 and 7 infrared radiometers (IRR) and proposal for Viking IRTM
- Anticipated need for global/annual models == big runs
- Computers were much slower than today
 - One case: 18 latitudes, 40 seasons, took an hour on the fastest unclassified computer in the world.
 - Now, one second
- Choose implicit forward differences
- Carefully code for speed
- Evolution: step-wise, with occasional bug fixes
 - Physics-based one-layer atmosphere
 - Additional output formats
 - Temperature-dependent thermal conductivity and specific heat



Major physics

- Upper and lower materials: separate K, ρ , C
- K and C may be temperature-dependent; cubic polynomial in T
- Terrain can slope, or be bottom of conical pit
- Insulating or fixed T lower boundary
- Any orbit, spin axis, rotation period.
- One layer atmosphere, gray in solar and thermal regions. Or none
- Atmosphere radiation, but not conduction.
- Conserves energy
- Local seasonal condensation. Atm. condensation falls as snow
- If surface reaches CO₂ saturation, starts frost accumulation
- Switch to frost properties; gradual for albedo
- Total frost may control global pressure, (the only latitude coupling)
or use Viking Lander relation, or constant
- Allows table-driven seasonal opacity and albedo
- Can use seasonal map of dust and ice opacity
- Version derived from Mike Smith THEMIS maps available
- Allows twilight

Basic Surface and Subsurface Treatment

- Surface condition: Radiation Balance
 - Specify albedo A , emissivity ϵ , slope and azimuth.
 - Can have frost with different A , ϵ and scattering thickness scale
 - Frost albedo: constant or linear with insolation (two precoded)
 - Imbalance of solar and thermal radiation goes into changing temperature of the top layer, or into changes of frost amount.
- Subsurface: thermal diffusion
 - Specify thermal inertia (I), conductivity(k) and density(ρ)
 - Lower boundary can be insulating, constant temperature, or fixed constant.
 - Can have a different material below a given layer;
 k , ρ and specific heat C_p .

Atmosphere and multiple scattering

- Gray in solar and thermal region, one layer,
 - Transmits, absorbs, and scatters solar radiation; Two stream Delta-Eddington
 - Transmits, absorbs and emits thermal radiation, Isotropic
 - Flux imbalance goes into change of temperature (apart from twilight)
- Dust amount and radiative characteristics can be specified
- Direct sunlight: Lambertian local surface -> DIRECT
- Diffuse Vis. & IR: based on fraction of sky visible -> DIFFUSE
- Reflected from surface -> BOUNCE
- All three Vis. above normalized to solar flux at Mars -> SOLR
- Surface pressure based on elevation. one of:
 - constant,
 - follow Viking Landers
 - based on global integration of frost.

2-stream Delta-Eddington for Sunlight

- ftp://climate1.gsfc.nasa.gov/wiscombe/Multiple_Scatt/DeltaEddington/DeltaEddingPaper.pdf
- Joseph et al, 1976 Delta-Eddington Approximation for radiative flux transfer, J.Atm.Sci vol.33,2453-2459 (1976)
- deding2.f See citations therein. Comments tie equations.
- Collimated and diffuse beams
- Scattering uses Henyey-Greestein function.
 - Assymetry set by ARC2 -> G0
 - Zero is isotropic, + 1 is pure forward scattering
 - Single scattering albedo set by DUSTA -> OMEGA . Maximum 0.999
- Runs in double precision (required)

Method

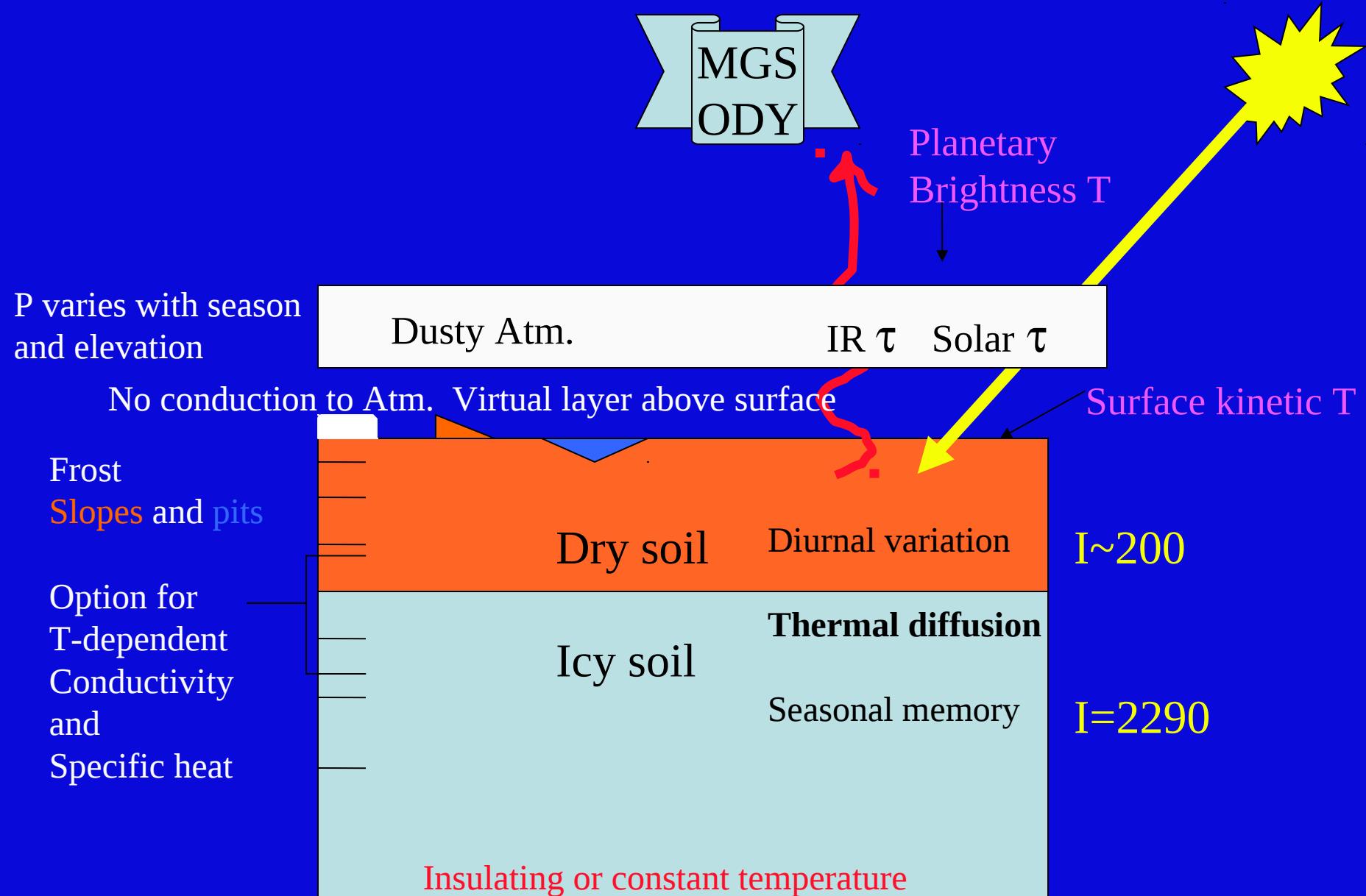
- Surface thermal inertia, specific heat and density determine the conductivity and diffusivity. Layers scaled with the diurnal skin depth (DSD)
 - First layer is 0.18 of DSD, successive ones thicker by factor of 1.2
 - As layers become so thick that thermal diffusion through them is small in one time step, time steps are doubled
 - Elaborate start-up and convergence techniques; Stability tests

This scheme allows accurate practical treatment of diurnal and seasonal heat flow

Virtually all physical and grid parameters adjustable: ~127 total.

Most you need not touch.

Cartoon of KRC model



What is in KRC

- Periodic insolation (sunlight)
 - Brightness of the Sun (or any star)
 - Orbit of mars (or any other elliptical orbit)
 - Orientation of Mars spin axis (or any planet/satellite/asteroid)
 - Length of a “day” (Sun on the meridian)
 - Any number of seasons/year, up to many years
- Up to 37 latitudes in a set. All with the same properties
- Up to two layered materials; any total depth
 - Density (constant), thermal conductivity, specific heat (may vary with T)
- Surface albedo; flat, plane slope, or conical depression
- Second-order dusty/icy atmosphere, or no atmosphere
- What's not included

Absent Geophysics. To be included?

- Radiative-convective atmosphere (Jakosky-Mellon)
- Turbulent mixing (Need wind speed; Vasavada; GCM)
- More than 2 materials; fully-flexible layer structure
- Vertical transport of sensible and latent heat by gaseous diffusion (Mellon, others) and regolith breathing (no one?)
- Local rocks and shadows: *lateral heat conduction*
- Regional topography with *shadows and radiative coupling*
- Season changes incrementally every SOL
 - Now: Planet spins in place for a season
- Geothermal heat flow
- Solar eclipse (for satellites)
- Equation of time (length of solar day varies through a year)

Atmosphere model.

- One Layer, gray in solar and thermal.
- Delta-Eddington, two-stream
- Starts at equilibrium with surface based on diurnal average radiation balance.
- Solar heating plus any infrared excess or deficit: change the temperature
- Affects view-from-space bolometric temperature

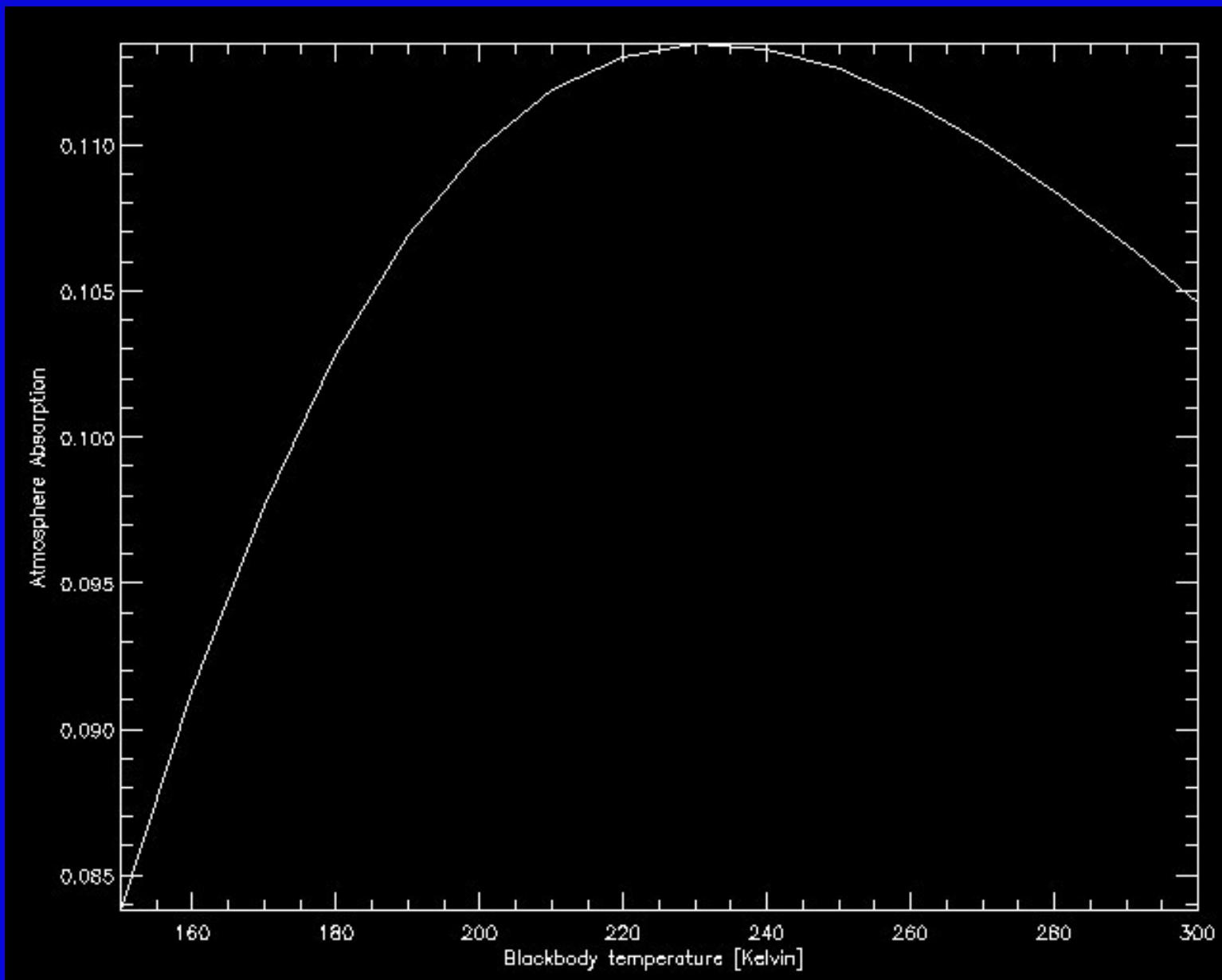
Slopes and pits

- Considers direct, diffuse, and one-bounce radiation.
- Approximation: regional surrounds at same temperature as target.

KRC atmosphere

- One layer; solar and thermal radiation (later slide)
- Gravity and molecular weight: default to Mars
- Set the scale height
$$P \propto e^{-z/\mathcal{H}} \text{ AND } \mathcal{H} = T_a \mathcal{R} / Mg$$
- Surface pressure: 3 options
 - Constant, \propto Viking, decreased by global frost mass.
- Dust opacity may vary with season: scaled to local pressure
- Global table, or zonal climate that allows ice-cloud opacity
 - Ice-cloud opacity does not scale to local surface pressure
- One-component may condense: snow and accumulated mass
- Presence of any frost sets surface T at the local saturation value
 - Frost albedo may vary with insolation intensity (Mars' does)
 - Frost effect on albedo goes realistically to zero as it thins

CO₂ gas only thermal radiation absorption



Energy Balance & Heat Diffusion Equations

Surface energy balance; excess/deficit into frost

$$W = (1 - A)S'_{(t)} + \Omega\epsilon R_{\downarrow} + k \frac{\partial T}{\partial z}_{(z=0)} - \Omega\epsilon\sigma T^4$$

Subsurface:

$$\frac{\partial T}{\partial t} = \frac{-1}{\rho c} \frac{\partial}{\partial z} \left(-k \frac{\partial T}{\partial z} \right) = \frac{k}{\rho c} \frac{\partial^2 T}{\partial z^2} + \left[\frac{1}{\rho c} \frac{\partial k}{\partial z} \frac{\partial T}{\partial z} \right]$$

Last term
assumed
zero

Finite
difference
code for grid

$$\frac{\Delta T_i}{\Delta t} = \frac{2}{B_i \rho_i C_i} \left[\frac{T_+ - T_i}{\frac{B_i}{k_i} + \frac{B_+}{k_+}} - \frac{T_i - T_-}{\frac{B_i}{k_i} + \frac{B_-}{k_-}} \right]$$

T-dependent
properties

Coded for T-ind as: DO J=2,KN with KN a “comb” for time-steps

DTJ(J)=FA1(J)*(TTJ(J+1)+FA2(J)*TTJ(J)+FA3(J)*TTJ(J-1))

Much more code for the T-dep case.

Followed by: DO J=2,KN

TTJ(J)=TTJ(J) + DTJ(J)

Equations accomodating T-dependence

[82] For variable layer thickness: equation (18) becomes

$$\frac{\Delta T_i}{\Delta t_i} = \frac{2}{B_i \rho_i C_i} \left[\frac{T_+ - T_i}{\frac{B_i}{k_i} + \frac{B_+}{k_+}} - \frac{T_i - T_-}{\frac{B_i}{k_i} + \frac{B_-}{k_-}} \right] \quad (19)$$

[83] For KRC, formulate this as

$$\Delta T_i = F_{1i}[T_+ + F_{2i}T_i + F_{3i}T_-] \quad (20)$$

[84] KRC define intermediate constants for each layer:

$$F_{1i} = \frac{2\Delta t_i}{B_i \rho_i C_i} \cdot \frac{1}{\frac{B_i}{k_i} + \frac{B_+}{k_+}} \equiv \frac{2\Delta t_i}{\rho_i C_i B_i^2} \cdot \frac{k_i}{1 + \frac{B_+}{B_i} \frac{k_i}{k_+}} \quad (21)$$

and

$$F_{3i} = \left(\frac{B_i}{k_i} + \frac{B_+}{k_+} \right) \cdot \frac{1}{\frac{B_i}{k_i} + \frac{B_-}{k_-}} \equiv \frac{1 + \frac{B_+}{B_i} \frac{k_i}{k_+}}{1 + \frac{B_-}{B_i} \frac{k_i}{k_-}} \quad (22)$$

and

$$F_{2i} = -(1 + F_{3i}) \quad (23)$$

For T-dependence,
must evaluate all k
for each time-step.

For T-ind, in each
equation the B
and k terms
reduce to a single
constant value for
each layer

More Equations

Convergence criterion

$$\frac{\Delta t}{(\Delta Z)^2} \kappa < \frac{1}{2} \quad \text{or} \quad (\Delta z)^2 \equiv B^2 > 2\Delta t \kappa$$

**Two-stream
Delta-Eddington**

$$F = \pi * [I_0 \pm \frac{2}{3} I_1]$$

where + is down, F^{\downarrow} ; - is up, F^{\uparrow}

**Atmospheric
heating**

$$H_V = S_M * \left(\mu_0 - F^{\uparrow}(0) - (1 - A_s) [\mu_0 \text{COLL} + F^{\downarrow}(\tau)] \right)$$

$$H_R = \epsilon \sigma T_s^4 (1. - e^{-\tau_e}) - 2R_{\downarrow t} = \sigma \beta_e (\epsilon T_s^4 - 2T_a^4)$$

$$\frac{\partial T_a}{\partial t} = \frac{H_R + H_V}{c_p M_a}$$

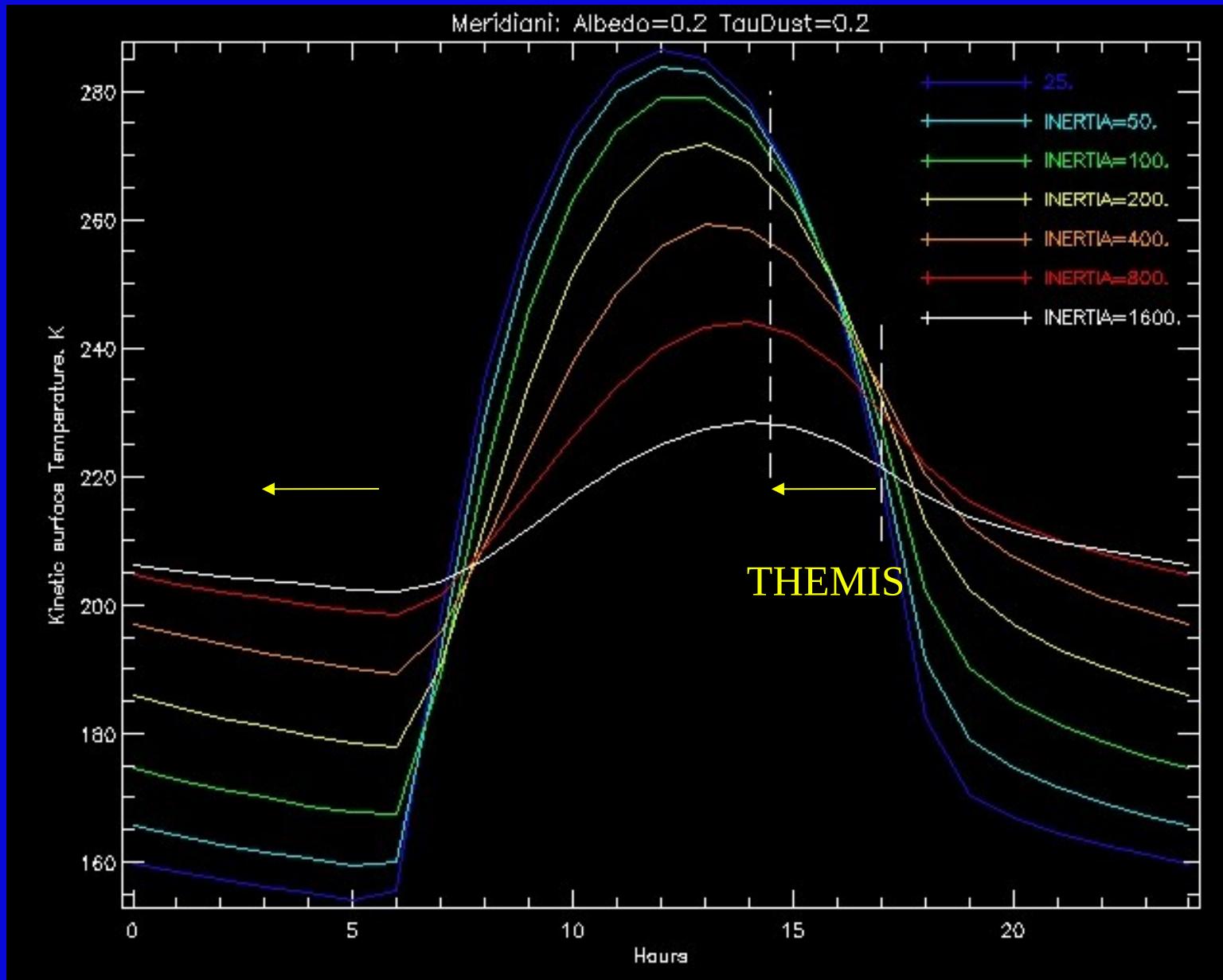
Planetary Temperature

$$\sigma T_P^4 = \epsilon \sigma T_S^4 (e^{-\tau_R}) + \sigma T_a^4 (1 - e^{-\tau_R}) \quad \text{or} \quad T_P = [\epsilon(1 - \beta) T_S^4 + \beta T_a^4]^{1/4}$$

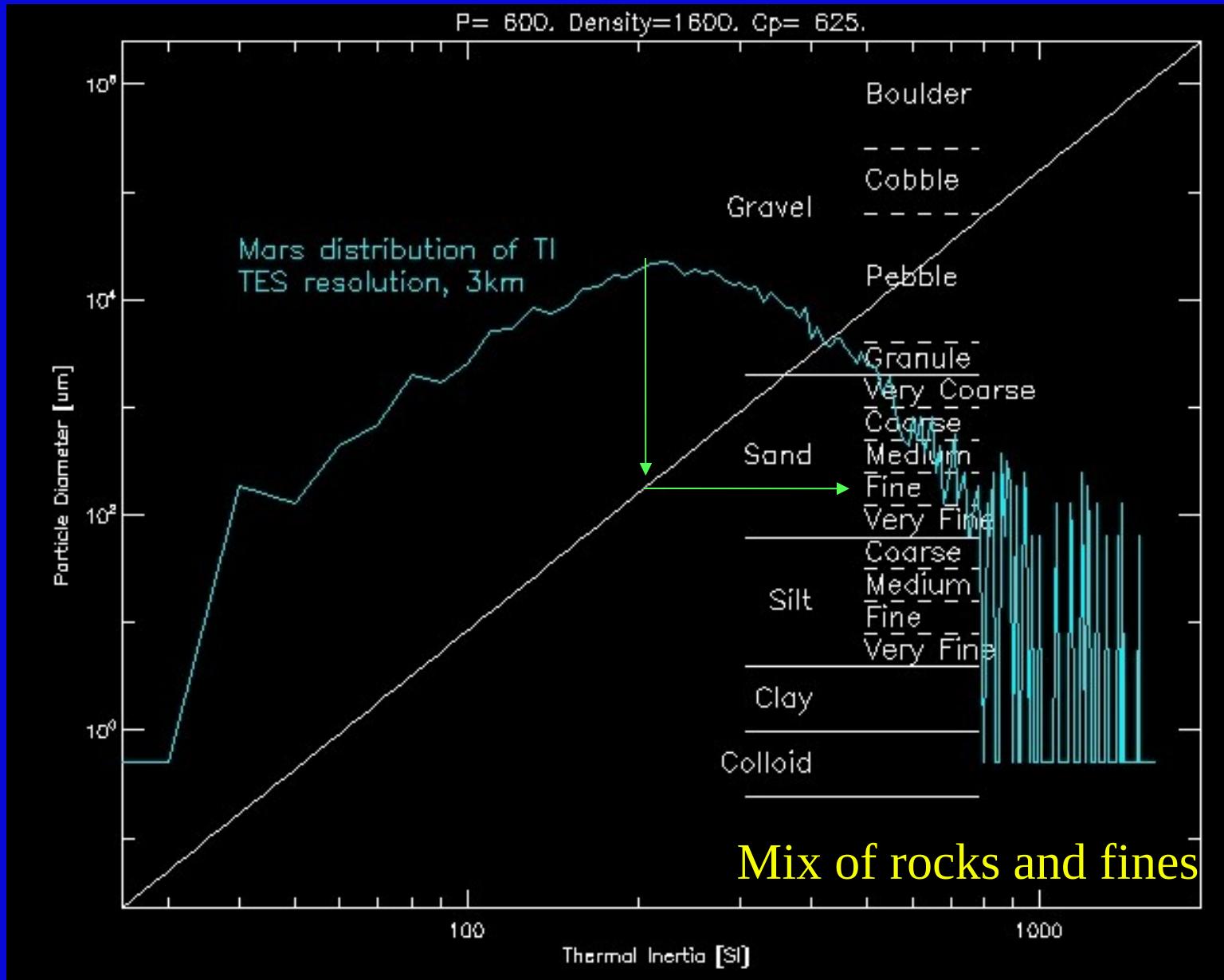
Numerical Approach

- Explicit forward finite differences
- Highly optimized code
- *Basic loop architecture*
- [Cases]
- Seasons
- Latitudes
- Days
- Layers-1
- Layers-2
- Up to 30 layers, up to 37 latitudes, up to 384*4 time-steps per sol, output at 24 or 48 “hours”/day, up to 16 days to converge., up to 740 seasons.
 - All the maximum sizes could be adjusted in krccom.inc
- Dynamic dimensioning of .bin5 output files [not FORTRAN-friendly]:
e.g., latitudes*seasons*cases

Thermal inertia effect on temperature



Effective particle size on Mars



Frost and snow

- Local seasonal CO₂ saturation at surface
- Super-cooling of atmosphere avoided: energy supplied by CO₂ condensation.
- Resulting snow falls to surface (within a day) and adds to surface frost inventory.
- Surface frost albedo can follow observed insolation-dependence
 - Mars' seasonal frost albedo commonly increases with Sun intensity
- Thin frosts are translucent, no discontinuities !
- Global integration of solid CO₂ :
 - average mass deducted from surface pressure

How close is KRC to Mars ???

- KRC based on best-estimates of physical parameters
- Absence of wind-effect probably the largest shortfall.
- Beware working near the edge of the seasonal cap:
 - Interpolation commonly not valid between frost and no-frost

Everything controlled by input file

- # layers, Hours_out/day, latitudes, seasons,...
- Start date, Delta season, first to disk, index of major changes
- Surface physical properties: two materials
 - neither/both use T-dependent properties
- Atmosphere properties
- Frost physical and radiative prop.
 - Option: insolation dependant albedo
- Planet orientation, day and orbit
 - Pre-computer rotation matrices
- Layer and time scheme; Convergence criteria
- Print output
- Output files: 3 direct-read (FORTRAN sequential access)
 - and 5 compound .bin5
 - Most users will be fine with Type 52

Meant to be a tool: Many ways to use it

- Sets of uniform global models
 - E.g., Viking used three
 - THEMIS using hundreds
 - Global average temperature map used nearly a thousand
- Studies of slope effect
- Models of rock proportion
- Multi-year “spin-up”, then every sol, hour or time step (5 minutes)
- The “one-point” mode:
- Input list of targets:
- Phil ran 1.5 billion points to make thermal movie
- There is a “Helplist” users guide
- KRC website provides ability to run the maintained version
- A distribution includes all source code, sample files, documentation,...

What do You want to do?

Several output formats: e.g. Type 52 has:

- Surface kinetic, planetary and atmosphere temperatures every “Hour”,
- minimum and maximum diurnal temperatures for each layer,

For each latitude:

Days to convergence, RMS temperature change on last day,
predicted atmosphere temperature, amount of frost, frost albedo,
mean upward heat flow into surface

Global values for every season:

Julian date (offset from 2,440,000), Ls, Atm pressure at 0 elevation,
dust opacity, total frost.

IDL readers provided for every output format.

Need to generate a common-language (non-IDL) reader

There is a DeVinci input interface to prior KRC version; could easily be updated

Input File: it's simpler than it looks

master.inp 2013jan22

0 0 / KOLD: season to start with; KEEP: continue saving data in same disk file

Default values for all parameters. 19 latitudes with mean Mars elevations

| | | | | | | | | | |
|----------|-----------|----------|-----------|----------|-----------|-----------|-----------|-------|-------|
| ALBEDO | EMISS | INERTIA | COND2 | DENS2 | PERIOD | SPEC_HEAT | DENSITY | | |
| .25 | 1.00 | 200.0 | 2.77 | 928.0 | 1.0275 | 647. | 1600. | | |
| CABR | AMW | [ABRPHA | PTOTAL | FANON | TATM | TDEEP | SpHeat2 | | |
| 0.11 | 43.5 | -0.00 | 546.0 | .055 | 200. | 180.0 | 1711. | | |
| TAUD | DUSTA | TAURAT | TWILII | ACR2 | [ARC3 | SLOPE | SLOAZI | | |
| 0.3 | .90 | 0.5 | 0.0 | 0.5 | -0.00 | 0.0 | 90. | | |
| TFROST | CFROST | AFROST | FEMIS | AF1 | AF2 | FROEXT | [FD32 | | |
| 146.0 | 589944. | .65 | 0.95 | 0.54 | 0.0009 | 50. | 0.0 | | |
| RLAY | FLAY | CONVF | DEPTH | DRSET | DDT | GGT | DTMAX | | |
| 1.2000 | .1800 | 2.0000 | 0.0 | 0.0 | .0020 | 0.1 | 0.1 | | |
| DJUL | DELJUL | SOLARDEC | DAU | Lsubs | SOLCON | GRAV | AtmCp | | |
| 10322.33 | 17.1745 | 00.0 | 1.465 | .0 | 1368. | 3.727 | 735.9 | | |
| ConUp0 | ConUp1 | ConUp2 | ConUp3 | ConLo0 | ConLo1 | ConLo2 | ConLo3 | | |
| 0.050087 | -0.002145 | 0.002347 | -0.000750 | 2.766722 | -1.298966 | 0.629224 | -0.527291 | | |
| SphUp0 | SphUp1 | SphUp2 | SphUp3 | SphLo0 | SphLo1 | SphLo2 | SphLo3 | | |
| 646.6275 | 246.6678 | -49.8216 | 7.9520 | 1710.648 | 721.8740 | 57.44873 | 24.37532 | | |
| N1 | N2 | N3 | N4 | N5 | N24 | IB | IC | | |
| 20 | 384 | 15 | 19 | 120 | 24 | 0 | 7 | | |
| NRSET | NMHA | NRUN | JDISK | IDOWN | FlxP14 | FlxP15 | KPREF | | |
| 3 | 24 | 1 | 81 | 0 | 45 | 65 | 1 | | |
| K4OUT | JBARE | Notif | [IDISK2 | | | | end | | |
| 52 | 0 | 20 | -0 | | | | 0 | | |
| LP1 | LP2 | LP3 | LP4 | LP5 | LP6 | LPGLOB | LVFA | LVFT | LkofT |
| F | T | F | F | F | F | F | F | F | F |
| LPORB | LKEY | LSC | spare | LOCAL | Prt76 | LPTAVE | Prt78 | Prt79 | L_ONE |
| T | @ASUF | F | F | T | F | F | F | F | F |

Input file, continued

Latitudes: in 10F7.2 7 7 7 7 7 7 7 7 7

-87.50 -80.00 -70.00 -60.00 -50.00 -40.00 -30.00 -20.00 -10.00 0.00

10.00 20.00 30.00 40.00 50.00 60.00 70.00 80.00 87.50 -0.00

Elevations: in 10F7.2 7 7 7 7 7 7 7 7 7

3.51 2.01 1.39 1.22 0.38 0.48 1.17 1.67 1.26 0.17

-0.94 -1.28 -1.99 -2.51 -3.52 -4.08 -4.51 -4.38 -2.57 -0.00

08 Sep 29 10:41:33 =RUNTIME. IPLAN AND TC= 4.0 0.55000

4.000000 0.5500000 0.8650615 0.3229325E-01 5.000821

0.9340634E-01 1.523671 12882.95 686.9650 0.9229904

5.544495 24.62280 0.000000 0.4093198 0.000000

0.000000 0.000000 0.000000 6.159676 0.4662921

0.4172604E-01 0.6197483 4.381073 0.000000 1.228627

0.6619807 0.000000 1.391099 0.1075499 -0.3195100E-01

0.2263214 -1.246176 -0.5861457 -0.8611114E-01 0.8908045

0.4461527 -0.9063585 0.1158813 -0.4063075 -0.4136413

-0.4393618 0.7974096 0.9138050 -0.4049719 -0.3095386E-01

0.4054090 0.9140893 0.9184200E-02 0.2457525E-01 -0.2094154E-01

0.9994786 -0.3252879 -0.8556869 -0.4024770 0.9456150

-0.2943530 -0.1384504 0.7823110E-07 0.4256245 -0.9048999

8 0 0 '/work1/krc/mars/masterA.t52' / Disk file name

1 12 540. 'PTOTAL set to yield 7 mb at VL1 @ Ls=100' /

1 4 0.22 'COND2' / to match lower k(T)

1 35 4. 'CONVF' / push time doubling start deeper

0/ a zero ends a case

3 10 1 'LkofT' / Temperature-dependant conductivity

0/

0/ two zeros end a run

Locations

Orbit
and
Pole
matrix.

Mars
2005

Change Cards,
See HelpList

Input Changes

2 integers, 1 real/integer/logical, one text [/] free comments

```
8 0 0 '/work1/krc/mars/master.t52' / Disk file name
1 12 540. 'PTOTAL set to yield 7 mb at VL1 @ Ls=100'
1 35 4. 'CONVF' / push time doubling start deeper
0/ <<< end of a case
0/ <<< 2nd consecutive 0 is end of the run
```

Next is an example of an elaborate set of change cards that looks in detail at the temperatures through the first 40 sols of ice freshly exposed at the bottom of a conical pit. It uses 3 latitudes and does 5 cases; the first is ice freshly exposed to a full hemisphere of sky, followed by pits with slopes of 45 and 25 degrees, then these two pits with a different initial ice temperature

Complex Input

```
7 7 7 'Pit dug to ice by Phoenix' / New title
8 0 0 '../output/phx4.t52' / Disk file name/
1 1 .20 'Albedo'
1 3 2025.3 'Inertia for ice' /
1 7 1300. 'Spec heat' / for ice
1 8 928. 'Density' / for ice
1 15 185. 'TDEEP' /
1 17 0.2 'TAUD'
1 39 .001 'GGT: set to avoid ending early' / set for daily output
1 41 11920.2 'DJUL' / starting date
1 42 1.0275 'DELJUL 1 sol' / set for daily output
2 1 19 'Num Layers' /
2 3 1 'N3: set to run each day' / set for daily output
2 4 3 'N4' / number of latitudes
2 5 40 'N5' / total number of seasons = sols
2 7 2 'IB start all =TDEEP' /
2 12 1 'JDISK start immediately' /
2 17 52 'K40UT' /
4 77 77 'New Latitudes' / Must be N4 of them in 10F7.2
  65.00  70.00  72.00 -10.00   0.00  10.00  25.00  45.00  70.00  22.00
5 77 77 'New Elevations' / Must be N4 of them in 10F7.2
  -3.5   -3.5   -3.5   00.0   00.0   00.0   00.0   00.0   00.0   -3.1
0/
1 24 -400. 'Azimuth. Set flag to indicate a pit' /
1 23 45. 'Slope' / slope of pit wall
0/
1 23 65. 'Slope' /
0/
1 23 45. 'reset Slope' /
1 15 220. ' TDEEP' /
0/
1 23 65. 'Slope' /
0/
0/
```

Layer Table: I=200 Cp=630 Density=1600

Conductiv.= 3.968E-02 Dens*Cp= 1.008E+06 Diffu.= 3.937E-08 Scale= 3.335E-02

| LAYER | THICKNESS | | CENTER_DEPTH | | CONVERGENCE | | | | | | | |
|----------------------------------|-----------|--------|--------------|---------|-------------|--------|---|----|----|----|----|----|
| | scale | meter | scale | meter | kg/m^2 | factor | | | | | | |
| 1 | 0.1800 | 0.0060 | -0.0900 | -0.0030 | 0.000 | 0.000 | | | | | | |
| 2 | 0.2160 | 0.0072 | 0.1080 | 0.0036 | 11.527 | 2.851 | | | | | | |
| 3 | 0.2592 | 0.0086 | 0.3456 | 0.0115 | 25.359 | 2.053 | | | | | | |
| 4 | 0.3110 | 0.0104 | 0.6307 | 0.0210 | 41.958 | 2.956 | | | | | | |
| 5 | 0.3732 | 0.0124 | 0.9729 | 0.0324 | 61.877 | 2.129 | | | | | | |
| 6 | 0.4479 | 0.0149 | 1.3834 | 0.0461 | 85.779 | 3.065 | | | | | | |
| 7 | 0.5375 | 0.0179 | 1.8761 | 0.0626 | 114.462 | 2.207 | | | | | | |
| 8 | 0.6450 | 0.0215 | 2.4673 | 0.0823 | 148.882 | 3.178 | | | | | | |
| 9 | 0.7740 | 0.0258 | 3.1768 | 0.1060 | 190.185 | 2.288 | | | | | | |
| 10 | 0.9288 | 0.0310 | 4.0282 | 0.1344 | 239.749 | 3.295 | | | | | | |
| 11 | 1.1145 | 0.0372 | 5.0498 | 0.1684 | 299.226 | 2.372 | | | | | | |
| 12 | 1.3374 | 0.0446 | 6.2758 | 0.2093 | 370.598 | 3.416 | | | | | | |
| 13 | 1.6049 | 0.0535 | 7.7469 | 0.2584 | 456.244 | 2.460 | | | | | | |
| 14 | 1.9259 | 0.0642 | 9.5123 | 0.3173 | 559.020 | 3.542 | | | | | | |
| 15 | 2.3111 | 0.0771 | 11.6308 | 0.3879 | 682.351 | 2.550 | | | | | | |
| 16 | 2.7733 | 0.0925 | 14.1730 | 0.4727 | 830.348 | 3.672 | | | | | | |
| 17 | 3.3279 | 0.1110 | 17.2235 | 0.5745 | 1007.945 | 5.288 | | | | | | |
| 18 | 3.9935 | 0.1332 | 20.8843 | 0.6966 | 1221.061 | 7.615 | | | | | | |
| 19 | 4.7922 | 0.1598 | 25.2771 | 0.8431 | 1476.800 | 10.965 | | | | | | |
| 20 | 5.7506 | 0.1918 | 30.5485 | 1.0189 | 1783.687 | 15.790 | | | | | | |
| 21 | 6.9008 | 0.2302 | 36.8742 | 1.2299 | 2151.951 | 22.737 | | | | | | |
| 22 | 8.2809 | 0.2762 | 44.4651 | 1.4831 | 2593.868 | 32.742 | | | | | | |
| 23 | 9.9371 | 0.3314 | 53.5741 | 1.7869 | 3124.169 | 47.148 | | | | | | |
| 24 | 11.9245 | 0.3977 | 64.5049 | 2.1515 | 3760.530 | 67.893 | | | | | | |
| 25 | 14.3094 | 0.4773 | 77.6219 | 2.5890 | 4524.163 | 97.766 | | | | | | |
| Bottom layers for time doubling: | | | | 2 | 4 | 6 | 8 | 10 | 12 | 14 | 25 | 31 |

T-dependent Table

| ----- TYPE LOC VALUE ----- | | | | Parameter changes | | | |
|----------------------------------|---------------|------------------------|--------|-------------------------|----------------|----------|----|
| Changed>> 3 10 1.000 LkofT LKoft | | | | | | | |
| Case 1 | DTIME: | total, user, system= | | 0.1600 | 0.1530 | 0.0070 | |
| CCKU/L= | 1.3000000E-02 | -5.2780001E-03 | | 2.6159999E-03 | -1.1220000E-03 | 2.786833 | -1 |
| .264019 | 0.5252116 | -0.4637450 | | | | | |
| CCPU/L= | 630.0000 | 240.3250 | | -48.54044 | 7.747490 | 1710.648 | 7 |
| 21.8740 | 57.44873 | 24.37532 | | | | | |
| | T | cond_SpHeat_UPPER_iner | | cond_SpHeat__LOWER_iner | | | |
| 130.0 | 0.02069 | 368.74 | 110.48 | 4.68794 | 1089.73 | 2177.33 | |
| 140.0 | 0.01947 | 402.71 | 112.01 | 4.37162 | 1157.44 | 2166.92 | |
| 150.0 | 0.01836 | 435.33 | 113.09 | 4.08806 | 1225.13 | 2155.87 | |
| 160.0 | 0.01735 | 466.66 | 113.82 | 3.83449 | 1292.94 | 2144.95 | |
| 170.0 | 0.01643 | 496.73 | 114.28 | 3.60811 | 1361.03 | 2134.75 | |
| 180.0 | 0.01560 | 525.61 | 114.54 | 3.40615 | 1429.53 | 2125.70 | |
| 190.0 | 0.01485 | 553.32 | 114.66 | 3.22583 | 1498.60 | 2118.05 | |
| 200.0 | 0.01417 | 579.93 | 114.66 | 3.06436 | 1568.38 | 2111.88 | |
| 210.0 | 0.01356 | 605.47 | 114.59 | 2.91895 | 1639.01 | 2107.07 | |
| 220.0 | 0.01300 | 630.00 | 114.47 | 2.78683 | 1710.65 | 2103.34 | |
| 230.0 | 0.01250 | 653.55 | 114.32 | 2.66522 | 1783.43 | 2100.24 | |
| 240.0 | 0.01204 | 676.19 | 114.13 | 2.55133 | 1857.52 | 2097.12 | |
| 250.0 | 0.01162 | 697.94 | 113.92 | 2.44238 | 1933.04 | 2093.15 | |
| 260.0 | 0.01124 | 718.86 | 113.68 | 2.33558 | 2010.15 | 2087.30 | |
| 270.0 | 0.01087 | 739.00 | 113.39 | 2.22816 | 2088.99 | 2078.34 | |
| 280.0 | 0.01053 | 758.39 | 113.05 | 2.11733 | 2169.72 | 2064.76 | |
| 290.0 | 0.01020 | 777.10 | 112.63 | 2.00031 | 2252.47 | 2044.81 | |
| 300.0 | 0.00988 | 795.16 | 112.10 | 1.87432 | 2337.39 | 2016.33 | |
| 310.0 | 0.00955 | 812.62 | 111.44 | 1.73657 | 2424.64 | 1976.71 | |
| 320.0 | 0.00922 | 829.53 | 110.60 | 1.58428 | 2514.35 | 1922.66 | |
| | ConUp | ConLo | SphUp | SphLo | | | |
| Input | 0.03968 | 3.40000 | 630.00 | 1710.00 | | | |
| T-dep | 0.01300 | 2.78683 | 630.00 | 1710.65 | | | |

1Typical Mars Display of print options, 3 latitudes 40 seasons

RSE 2-2 13 Jan 27 06:36:01 PAGE= 5

[More--(60%)]

← Check for similarity

RUN-C

The one-point mode

- Input file points to ‘one.inp’ list of “targets”.
- Interactive, or provide a text-file of rows with
 - Ls = season
 - Latitude
 - Hour of day
 - Surface elevation
 - Albedo
 - Thermal Inertia
 - Solar opacity
 - Surface slope
 - Slope azimuth
- Any other non-default values can be specified in the formatted input parameter file
- Outputs:
 - Surface kinetic temperature
 - Top-of-atmosphere bolometric brightness temperature

KRC-lite. The “one-point” interface

- You build a table of “points” of interest:
 - L_s
 - Latitude
 - Hour
 - Elevation (km)
 - Bolometric Albedo
 - Thermal Inertia (SI units)
 - Atmosphere Visual opacity
 - Regional slope (degrees)
 - Azimuth of down-slope (degrees Eastward of North)
- Output file has these, plus columns of:
 - Surface kinetic temperature, T_{surf}
 - Top-of-Atmosphere bolometric temperature, T_{plan}=T_{bolo}

One-point details

- Copy one.inp Edit it; rigid format. Keep name the same
 - Title Line, Text transferred to output. This & next line must be present
 - 11 Ls Lat Hour Elev Alb Inerti Opac Slop_Azim <Align each field
 - 11 123.1 4.1 9.4 22.22 0.17 100.0 0.10 0 0
 - 11 320.1 -33.1 15.5 -3.12 0.15 500.1 0.55 0.0 123.
 - 11 180. -10.1 13.5 0. 0.20 200.0 0.50 0.0 123.
- Run krc. Input file is: krcone_master.inp
 - This file sets KRC into 13 layer, 1 latitude, short season mode, and points to one.inp.
- Underneath the one-point mode is the full KRC model
- You can copy krcone_master.inp and edit it to do [almost] whatever you like, including point to a different one-point file

Production Mode

- Pre-compute hundreds of models on a grid of albedo, dust opacity, thermal inertia, surface elevation, latitude, hour, season.
Traditionally output direct-access files, a record for each season
- Interpolation scheme: mostly quadratic
- Extract surface and nadir-view temperatures for target.
- Select results for albedo and dust opacity based on observations
- Logarithmic interpolate at observed temperature to derive thermal inertia
- Time is not an issue. Can do entire TES or THEMIS model set* in an hour.
 - Mellon models took a week.
- * 48 times of day, 37 latitudes, [168+] 84 seasons, 3 albedos, 3 opacities, 3 elevations, 10 inertias.

Output

- Default print file (`krc.prt`) lists all changes, then the full input list, then the layer table, which should usually be contemplated.
- Options to turn on many kinds of printout.
 - Beware: file can get huge.
- Binary output files: 7 kinds in two sets: controlled by `K4OUT`
- Fortran direct access. [Allows restart at any recorded season]
- -1: Used for major model runs. Outputs `KRCCOM`, which contains all the input parameters, then arrays of Surface kinetic temperature and Planetary Bolometric Temperature for each season.
- 0: `LATCOM`. Arrays of several variables versus latitude and depth
- +1: `DAYCOM` for the last latitude for each season. Arrays of variables versus time-step or Hour or convergence-day.

Bin5 Binary output arrays

- +51: Minimum and Maximum temperature for all layers, 5 latitudes, 40 seasons, 5 cases
 - Layers 29,30 contain Tp_min, Tp_max, L_sub_s, Julian Day
- +52: 24 hours, 6 items, 5:10 lats, 1+<=80 seasons, <=5 cases)
 - Items are: Tsurf, Tplan, Tatm, DownVis, DownIR, spare
 - First "season" contains: KRCCOM and 3 season variables
- +53: Combo at 1 lat, 2+80 seasons, 10 cases
 - Combo: Tsurf & Tplan @ 24 hours; Tmin & Tmax @ 30 layers,
 - Asol @ 24 hours. The first 2 "seasons" contain KRCCOM
- +54: 68:740 seasons, 3 items, 1+<= 5 lats, <=10 cases
 - 3 items are: heat flow, surface temperature, frost amount
 - Store KRCCOM in the first latitude

Model sets and Determination of Thermal Inertia

- Two large set of models have been pre-computed.
 - Both have: Two year run-up, 3 albedos, 3 opacities, 3 elevations, logarithmic inertia. Tsur and Tplan output at 24 Hour points:
- Mellon: Used for TES production.
 - 84 seasons, 39 latitudes, 10 Inertias (25:800)
- KRC: recent run, plan to extend to $I \geq 2000$
 - 84 seasons, 37 latitudes, 15 Inertias (25:800), 3 albedos, 3 pressures, climate odust and ice opacities, 9 slopes [K4OUT=-1]
- Interpolation: described in Mellon,Jakosky, Kieffer and Christensen (2000): Icarus, 148, on page 441.4a
 - Season: linear
 - Albedo, opacity, elevation: quadratic
 - Hour and Latitude: bilinear
- Yields a set of 10 temperatures. Finds set between T_{min} and T_{max}
 - Inertia: linear in logarithm, And dT/dI for final interval

Installation and Supporting documentation

KRC installed at ASU

Helplist.tex Guide to running KRC

/mars/u/hkieffer/cop/krcCorP2.pdf Corrected version of JGR article

HeatOft.pdf Temperature dependence of rock thermal properties

Cement.pdf Particulate thermal conduction

newKRC.ppt 2008 powerpoint with some more details

mkrc.pdf making thermal movies

Have a look at ~/phil/ls09_32_full_5_part_movie.mov

Expect all this to be in Team-access area.

Will become public upon publication of the KRC paper

Coding Philosophy

- Place the large-loop actions (seasons, latitude) in separate routines
- Communicate via commons
- All dimensions set in the main common: krccom
- Minimize operations in the inner-most loop
- Implicit none in all routines that use commons
- tday.f: top section sets up the layer and time system
- Called once per case

A few aspects of FORTRAN

One of the earliest languages (FORmula TRANslator) and still widely used for heavy-duty computing (large scientific models)
Language has evolved, but largely backward compatible.

1-based indexing

Default typing based on first letter of name

I,J,K,L,M,N: default-length (4-byte) integers

A-H,O-Z: single-precision floating-point (REAL)

However, KRC mostly uses IMPLICIT NONE (contained in krccom.inc), which requires that each variable be explicitly typed.

Subroutine values NOT remembered by default, but can be SAVED or placed in (named) COMMON block

EQUIVALENCE puts items in the *same place in memory*, including all related to it

Some bad habits (speedy treachery) in KRC

Pitiful excuse: 8 Mbyte took half a floor!

Overloading: use the same variable for different things at different times

Equivalence: can place different things in the same place in memory

Use of REAL vectors to hold strings

DO LOOP indices placed in COMMON.

Allows reporting just where program was at an error

Use of COMMON simply to ensure remembrance

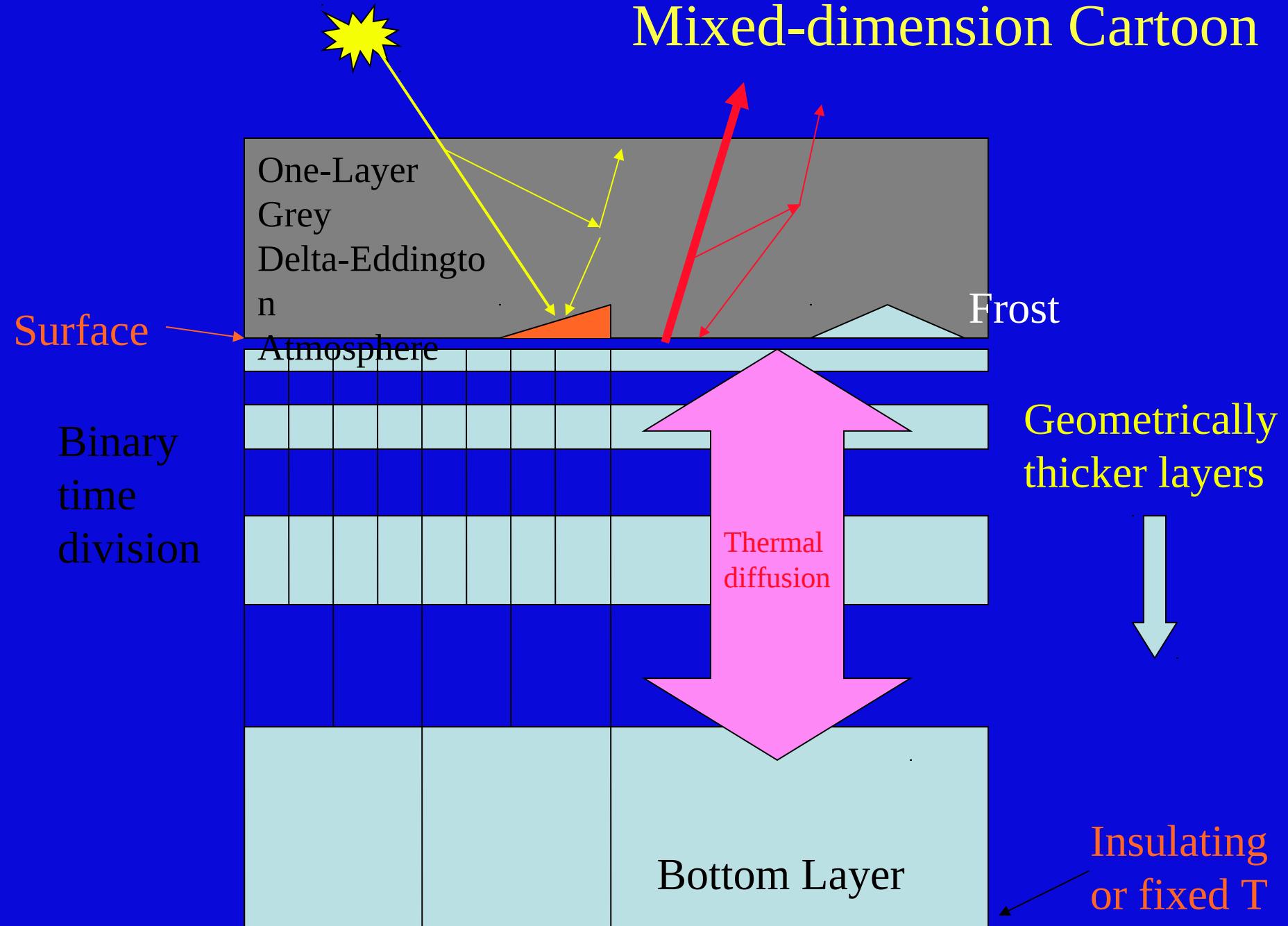
Use of subroutines to move things in memory.

E.g., r2r: (from,to,number(of 4-byte words)) N negative \Rightarrow from is a constant

Does not care about actual types ! A seductive tool.

The C version checks for physical location, and runs backwards if needed.

Mixed-dimension Cartoon



Flow: Loops and Calls

\sim : optional

. . : only if atmosphere

K: only if T-dependent

Latitude Common: latcom.inc

```
C_Titl LATCOM.INC common for latitude-dependant items in KRC
      INTEGER NWLAT ! size of common in 4-byte words
      PARAMETER (NWLAT= (9+ 3*MAXN1 + 2*MAXNH) *MAXN4) !
C_arg: all are set in TLATS
      INTEGER NDJ4(MAXN4) ! # days to compute solution for each latitude
      REAL DTM4(MAXN4)    ! rms temperature change on last day
      8,TST4(MAXN4)       ! predicted equilibrium temperature of ground
      8,TTS4(MAXN4)       ! predicted mean surface temperature for each latitude
      8,TTB4(MAXN4)       ! predicted mean bottom temperature
      8,FROST4(MAXN4)    ! predicted frost amount kg/m^2.
      8,AFR04(MAXN4)     ! frost albedo.
      8,TTA4(MAXN4)       ! predicted final atmosphere temperature
      8,TTX4(MAXN4)       ! spare
      8,TMN4(MAXN1,MAXN4) ! predicted convergence midnight temperature
      8,TIN(MAXN1,MAXN4)  ! minimum hourly layer temperature
      8,TAX(MAXN1,MAXN4)  ! maximum hourly layer temperature
      8,TSF(MAXNH,MAXN4) ! final hourly surface temperature
      8,TPF(MAXNH,MAXN4) ! final hourly planetary temperature
C_Var: all are set in TLATS
      COMMON /LATCOM/ NDJ4,DTM4,TST4,TTS4,TTB4,FROST4,AFR04,TTA4,TTX4
      8,TMN4,TIN,TAX,TSF,TPF
```

C_Desc

krccom.inc top part

```
C_Titl KRCCOM.INC common for input and transfer variables
C_Limitations
    IMPLICIT NONE ! none-the-less, try to code with usage
C IMPLICIT REAL (A-H,O-Z), INTEGER (I,J,K,M,N), LOGICAL*4 (L) !std.,+L
    INTEGER MAXN1,MAXN2,MAXN3,MAXN4,MAXN5,MAXN6,MAXNH,MAXBOT
    8,MAXN1P,NUMFD,NUMID,NUMLD,NWKRC,KOMMON
C Here are all the dimension-defining parameters for items in any common
    PARAMETER (MAXN1 =30)      ! dimension of layers
    PARAMETER (MAXN2 =384*4)    ! dimension of times of day: 384=24*16
    PARAMETER (MAXN3 =16)       ! dimension of iteration days
    PARAMETER (MAXN4 =37)       ! dimension of latitudes
    PARAMETER (MAXN5 =161)      ! dimension of saved seasons
    PARAMETER (MAXN6 =6)        ! dimension of saved years
    PARAMETER (MAXNH =48)       ! dimension of saved times of day
    PARAMETER (MAXBOT=6)        ! dimension of time doublings
    PARAMETER (MAXN1P=MAXN1+1)  ! dimension layer temperature points
    PARAMETER (NUMFD=64+32, NUMID=40, NUMLD=20)! number of each type
    PARAMETER (NWKRC=NUMFD+NUMID+NUMLD+25+2*MAXN4)
                                ! above is size of common in 32-bit words
C     PARAMETER (KOMMON=4*1056720) ! 4-byte words: Type 52: 48 hours, 37 lats
,84 seasons
C     PARAMETER (KOMMON=50000000) ! 4-byte words: Type 52: 34 hours, 36 lats,8
0 seasons 10 cases
C     PARAMETER (KOMMON=100000000) ! 4-byte words

    INTEGER N1,N2,N3,N4,N5, N24,IB,IC,NRSET,NMHA          ! 1:10
    8,NRUN,JDISK,DOWN,I14,I15, KREF,K4OUT,JBARE,NMOD,DISK2 ! 11:20
    8,KOLD,KYALB,KYTAU,LD24(4), NFD,NID,NLD              ! 21:30
    8,N1M1,NLW,JJO,KKK,N1PIB, NCASE,J2,J3,J4, J5        ! 31:40

    REAL ALB,EMIS,SKRC,COND2,DENS2, PERIOD,SPHT,DENS,CABR,AMH ! 1:10
    2,ABRPHA,PTOTAL,FANON,TATH,TDEEP, SPHT2,TAUD,DUSTA,TAURAT,TWILI ! 11:20
    3,ARC2,ARC3,SLOPE,SLOAZI,TFROST, CFROST,AFROST,FEMIS,AF1,AF2 ! :30
    4,FROEXT,FD32,RLAY,FLAY,CONVF, DEPTH,DRSET,DDT,GGT,DTMAX ! :40
    5,DJUL,DELJUL,SDEC,DAU,SUBS, SOLCON,GRAY,ATMCP ! :48
    8,HUGE,TINY,EXPMIN,FSPARE,FLOST,RGAS,TATHMIN,PRES,OPACITY,TAUR ! :74
```

-More--(36%)

krccom.inc Part 1

C Here are all the dimension-defining parameters for items in any common

```
PARAMETER (MAXN1 =30) ! dimension of layers
PARAMETER (MAXN2 =384*4) ! dimension of times of day: 384=24*16
PARAMETER (MAXN3 =16) ! dimension of iteration days
PARAMETER (MAXN4 =37) ! dimension of latitudes
PARAMETER (MAXN5 =161) ! dimension of saved seasons
PARAMETER (MAXN6 =6) ! dimension of saved years
PARAMETER (MAXNH =48) ! dimension of saved times of day
PARAMETER (MAXBOT=6) ! dimension of time doublings
PARAMETER (MAXN1P=MAXN1+1) ! dimension layer temperature points
PARAMETER (NUMFD=64+32, NUMID=40, NUMLD=20)! number of each type
PARAMETER (NWKRC=NUMFD+NUMID+NUMLD+25+2*MAXN4)
```

! above is size of common in 32-bit words

C PARAMETER (KOMMON=4*1056720) ! 4-byte words: Type 52: 48 hours, 37 lats,84 seasons
C PARAMETER (KOMMON=5000000) ! 4-byte words: Type 52: 34 hours, 36 lats,80 seasons 10 cases

PARAMETER (KOMMON=10000000) ! 4-byte words

```
INTEGER N1,N2,N3,N4,N5, N24,IB,IC,NRSET,NMHA                 ! 1:10
&,NRUN,JDISK,DOWN,I14,I15, KPREF,K4OUT,JBARE,NMOD,DISK2 ! 11:20
&,KOLD,KVALB,KVTAU, ID24(4), NFD,NID,NLD                 ! 21:30
&,N1M1,NLW,JJO,KKK,N1PIB, NCASE,J2,J3,J4,    J5         ! 31:40
```

Inclusion of commons

Copy from re-arranged Makefile

```
krc.o: krc.f krccom.inc latcom.inc daycom.inc units.inc filcom.inc  
tseas.o: tseas.f krccom.inc latcom.inc hatcom.inc units.inc  
tlats.o: tlats.f krccom.inc latcom.inc daycom.inc hatcom.inc units.inc  
mlats.o: mlats.f krccom.inc latcom.inc daycom.inc hatcom.inc units.inc  
tday.o: tday.f krccom.inc daycom.inc hatcom.inc units.inc  
tcard.o: tcard.f krccom.inc latcom.inc daycom.inc units.inc filcom.inc  
tprint.o: tprint.f krccom.inc latcom.inc daycom.inc units.inc filcom.inc  
tdisk.o: tdisk.f krccom.inc latcom.inc daycom.inc hatcom.inc units.inc filcom.inc  
tint.o: tint.f krccom.inc  
seasalb.o: seasalb.f units.inc filcom.inc  
seastau.o: seastau.f units.inc filcom.inc  
climtau.o: climtau.f units.inc filcom.inc  
porb0.o: porb0.f units.inc porbcm.inc  
albvar.o: albvar.f krccom.inc  
alsubs.o: alsubs.f krccom.inc
```

krccom.inc Lower part

COMMON /KRCCOM/

1 ALB,EMIS,SKRC,COND2,DENS2, PERIOD,SPHT,DENS,CABR,AMW ! 1:10
2,ABRPHA,PTOTAL,FANON,TATM,TDEEP, SPHT2,TAUD,DUSTA,TAURAT,TWILI ! 11:20
3,ARC2,ARC3,SLOPE,SLOAZI,TFROST, CFROST,AFROST,FEMIS,AF1,AF2 ! :30
4,FROEXT,FD32,RLAY,FLAY,CONVF, DEPTH,DRSET,DDT,GGT,DTMAX ! :40

Cset *day1
5,DJUL,DELJUL,SDEC,DAU,SUBS, SOLCON,GRAV,ATMCP ! :48

Cset v total of 56 input
5,CCKU,CCKL,CCPU,CCPL,HUGE,TINY ! :66
&,EXPMIN,FSPARE,FLOST,RGAS,TATMIN,PRES,OPACITY,TAUIR,TAUEFF,TATMJ ! :76

Cset *seas ----seas-----
6,SKYFAC,TFNOW,AFNOW,PZREF,SUMF, TEQUIL,TBLOW, HOURO,SCALEH,BETA ! :86

Cset -----lats----- tint ---lats----- ---lats---
6,DJU5,DAM,EFROST,DLAT,COND, DIFFU,SCALE,PIVAL,SIGSB,RADC ! :96

Cset seas lat day2 lats -----day1----- --main-----
7,N1,N2,N3,N4,N5, N24,IB,IC,NRSET,NMHA ! 1:10
8,NRUN,JDISK,IDOWN,I14,I15, KPREF,K4OUT,JBARE,NMOD,DISK2 ! 11:20
9,KOLD,KVALB,KVTAU,ID24, NFD,NID,NLD ! 21:30

Cset ----card---
A,N1M1,NLW,JJO,KKK,N1PIB, NCASE,J2,J3,J4, J5 ! 31:40

Cset ---day1- lat ----day1- main -day2- lats seas
B,LP1,LP2,LP3,LP4,LP5, LP6,LPGLOB,LVFA,LVFT,LKOFT ! 1:10
C,LPORB,LKEY,LSC,LNOTIF,LOCAL, LD16,LD17,LD18,LD19,LONE ! 11:20
D,TITLE,DAYTIM,ALAT,ELEV !

Cset tcards tprint tcards tcards

C

EQUIVALENCE (FD(1),ALB), (ID(1),N1), (LD(1),LP1)

What makes it fast

- Geometrically increasing layer thickness with depth
- Progressive time-step doubling as layer thickness allows for numerical stability
- Jump perturbations to equalize average temperature with depth in the first season
- Asymptotic predictor after convergence to the end of a season
- Pre-calculate frequently-used terms (Modern optimizer would do this)

Major Limitations

- Can not do really thin surface layers

Some Details

TBLOW: stops if this is exceeded by surface kinetic temperture

Twice the temperature of black plate facing the Sun

$$TBLOW=2.*\sqrt{\sqrt{[(\text{Solar constant}/\text{distance}^2)/(\text{Emissivity}*\sigma)]}}$$

Heliocentric distance, Solar declination, Ls

PORB returns, for a modified Julian Date (JD-2 440 000),

DAU=heliocentric distance in AU(distance for which Solar Constant is defined)

PEA,PEB=latitude,longitude(radians) of planet in heliocentric ecliptic coord.

HFA,HFB=latitude,longitude(radians) of Sun in planetocentric fixed coordinates

Packing of bin5 files

All done and remembered in TDISK (tdisk.f)

When called with a 1: sets up 10-element JJJ with sizes that will go into the bin5 header

- 1) the number of dimensions to the array
- 2:N+1) size of those dimensions

Type 52: (hours,7,latitudes,x+seasons,cases)

First x seasons contain, for each case:

Float of 4 integers that define sixes

Total size of krccom transferred

1-based index of the dimension with extra values

Number of those extra dimensions

spare

Floating point of krccom

Array(nseas,5) DJU5,SUBS,PZREF,TAUD,SUMF

For true seasons, the 7 contain:

1: TSF, surface kinetic temperature

2: TPF, planetary bolometric temperature

3: TAF. Atmosphere kinitec temperature

4: DOWNVIS:

5: DOWNIR

6: packed: float(NDJ4)+ DTM4 + TTA4+ Tmin(Nlay-) || Omitting the virtual

7: packed: FROST4+ AFRO4+ HEATMM+ Tmax(Nlay-) || first layer

Reading packed bin5 files

readk52.pro, finame,ttt,uuu,vvv,itemt,itemu,itemv, ddd,ggg,itemd,itemg ,t51=t51

```
;finame in. String of file name
; ttt  out.  Fltarr(hour,item,latitude,season,case) Item labels are in itemt
;          0= surface kinetic temperature
;          1= Top-of-atmosphere bolometric temperature
; t52    2= one-layer atmosphere kinetic temperature
; t52    3= Down-welling solar radiance
; t52    4= Down-welling thermal radiance
; uuu  out.  Fltarr(nlat,item,case) Item labels are in itemu
;          0= Latitude in degrees
;          1= elevation in Km.
; vvv  out.  Fltarr(season,item,case) Item labels are in itemv
;          0= Model season julian date - 2,440,000
;          1= L-sub-S computed in KRC
;          2= Global mean pressure: PZREF
; itemt,u,v  out.  Strarr ID's for the items in ttt, uuu, and vvv
; Following 4 outputs not valid for Type 51
; ddd  out.  Fltarr(layer,item,latitude,season,case) Item labels are in itemd
;          Item: 0=Tmin  1=Tmax
; ggg  out.  Fltarr(item,latitude,season,case) Item labels are in itemg
;          Items: 'NDJ4','DTM4','TTA4','FROST4','AFRO4','HEATMM'
; itemd,g out.  Strarr ID's for the items in ddd, and ggg
; t51  in_   Flag. If set, presumes input file is type 51
; func. out.  Fltarr of L-sub-S corresponding to Jul.Day for the first case.
;          If an error, returns negative integer
;      -1: -4 are from READKRCCOM, -5 = failure here
```

Readkrc52: lower calls

bin5.pro: BIN5,'R',finame,head,aaa,/verb

reads the bin5 packed array

readkrccom.pro:

;Use: Call initially, q=READKRCCOM('filename',hold) where hold may be undefined

; To get a case kcom=READKRCCOM(Icase,hold) where hold is unchanged

; When done with file: READKRCCOM(-1,hold) or free_lun,hold[0]

;_Calls DEFINEKRC

Hold is 4 words: [lun, size-of-case, # cases, nwkrc]

definekrc.pro: Define structures in IDL that correspond for Fortran commons

Basically a translation of krccom.inc, latcom, daycom or filcom into IDL, with default or fake values. Therefor MUST AGREE with the krccom used to write the file.

PORB: Planetary ORBit calculator

PORBMN Main program. Includes 'porbcm.inc'

DATIME Returns current date and time

may call the following in any order. All INCLUDE 'porbcm.inc'

PORB1 Read orbital elements, compute matrices

PORBEL Read any of 4 orbital element files, compute basic constants

YMD2JD Convert year, month, day to Julian date offset from 2,440,000

OBLIQ Compute Planets poles in ecliptic coord and nodes

OBLIP Compute obliquity of a planet

ROTDIA 2x

ROTAX 10x

PORBIO Read/write Common to file as text or binary

EPHEMR Print ephemeris == geometry versus time

PORB Computes planetary angles and location for specific time

ORBIT Compute location of body in its orbital plane

ROTVEC Apply rotation matrix to rotate a vector

CALDATE

SPCREV

ANG360

PORBQQ Test computation of matrices

ROTDIA 2x

ROTAX 4x

MPRINT 3x

MPROD3 2x

PORB: Planetary ORBit calculator

Runs as stand-alone program porbmn.f

Subroutine PORB used by KRC

Key input are the [four] elements formatted-text files

Users guide: PORBguide.txt

```
PORBMN Main program. Includes 'porbcm.inc'
DATIME Returns current date and time
may call the following in any order. All INCLUDE 'porbcm.inc'
PORB1 Read orbital elements, compute matrices
PORBEL Read any of 4 orbital element files, compute basic
YMD2JD Convert year, month, day to Julian date offset fr
OBLIQ Compute Planets poles in ecliptic coord and nodes
OBLIP Compute obliquity of a planet
ROTDIA 2x
ROTAX 10x
PORBIO Read/write Common to file as text or binary
EPHEMR Print ephemeris == geometry versus time
PORB Computes planetary angles and location for specific
ORBIT Compute location of body in its orbital plane
ROTVEC Apply rotation matrix to rotate a vector
CALDATE
SPCREV
ANG360
PORBQQ Test computation of matrices
ROTDIA 2x
ROTAX 4x
MPRINT 3x
MPROD3 2x
```

Debug print options

If first input line has third non-zero integer, then will read a an extra
2nd line of six integers: IDB1, IDB2, IDB3, IDB4, IDB5, IDB6

These control optional WRITE statements as listed in helplist section
on DEBUG OPTIONS

Restart for prior run. Powerful & delicate

First input line: KOLD & KKEEP [& KDB]

KOLD: record of prior run to start with. Will fill KRCCOM and Latitudes and elevations from prior run, and use the kinetic tempertures of the layers at the end of the specified season.

Must not change the layering scheme.

KKEEP: Flag, 0=do not write to this file. Else, leaves the file open, And will start writing based on JDISK

Not sure this still works:

If an error occurs:

Compare your input file to the master.

Unix: use diff

Check the helplist definition for each item that was modified

If can not get it to run properly ---

Send three files to the mentor:

Input file

What appeared on the screen

Print file

Liens

binf5.F Detecting hardware architecture. Near line 122.

Currently does not detect, so swap might no happen.

Bastion : uname -p = i686 -i = i386

Hulk2: x86_64 for both

Add version: to print, monitor and bin5. Here are all the steps:

transfer through filcom.inc: CHARACTER*12 VERDAT

krc.f VERDAT='KRC2013jan27'

 WRITE (IOPM,*)"This is:",VERDAT

tdisk.f HEADER --> HEADTX (header appears in many commens)

CHARACTER*3 CSTAT ! File status and string version of KODE

CHARACTER*25 HEADTX

write(cstat,*)kode ! converts file type to string

HEADTX=VERDAT//'tdisk Type'//CSTAT ! 12+10+3

tprint.f at 800, replace ' RUN-CASE' with VERDAT as A12

Liens: 2

Decide how to handle “master” version of code during transition

Automate processing of a new Distribution

Helplist:

- Rename .tex to .txt

- Make html version with index

- Make pdf version with index

Remove the alternate input options (not used for a long time)

?? Remove the debug options

Resolve the extra-word issue between FORTRAN and IDL

Make a re-ordered entry form for input: output the current input file

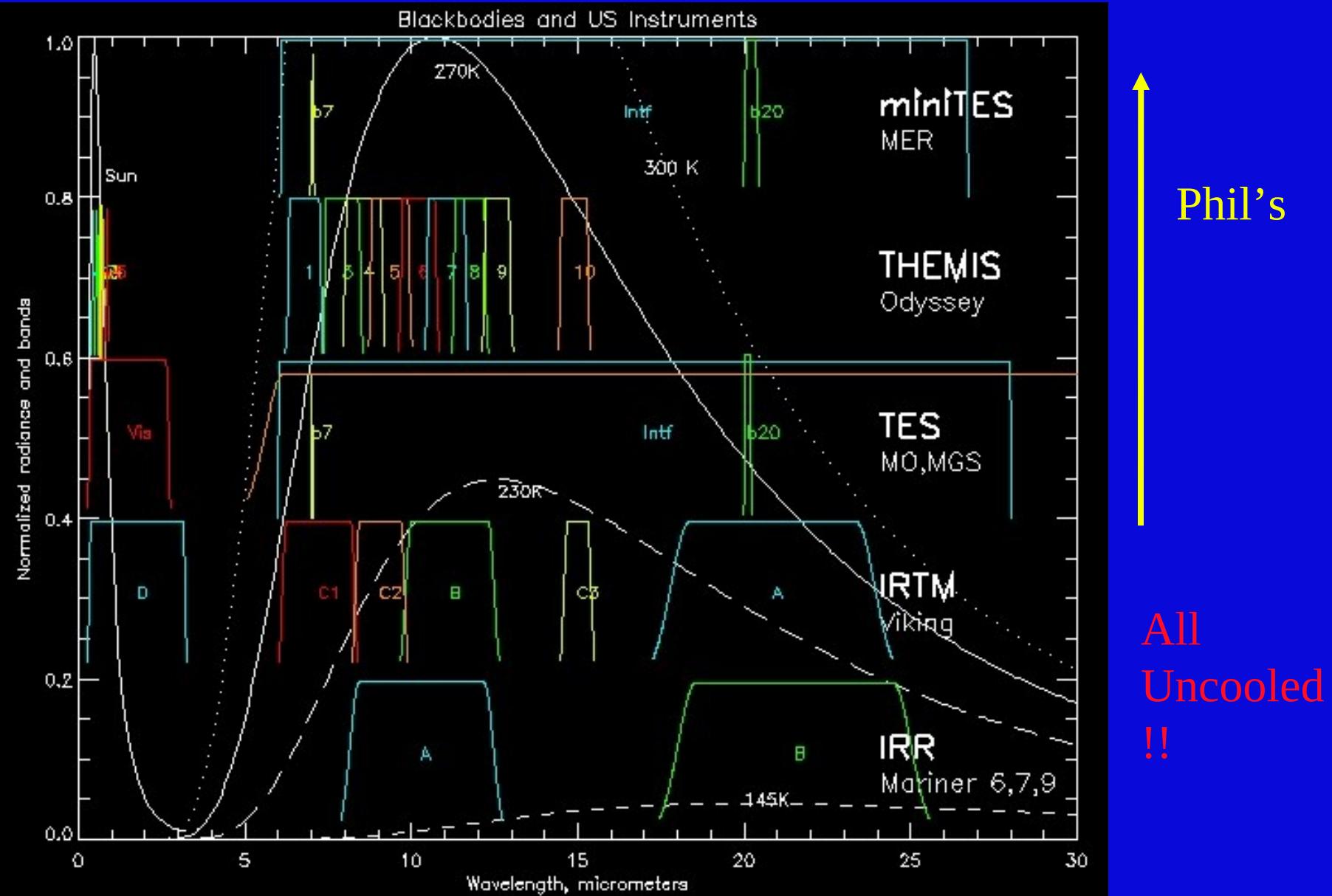
- Real time limit checks (some are linked)

- Put expert-only items last

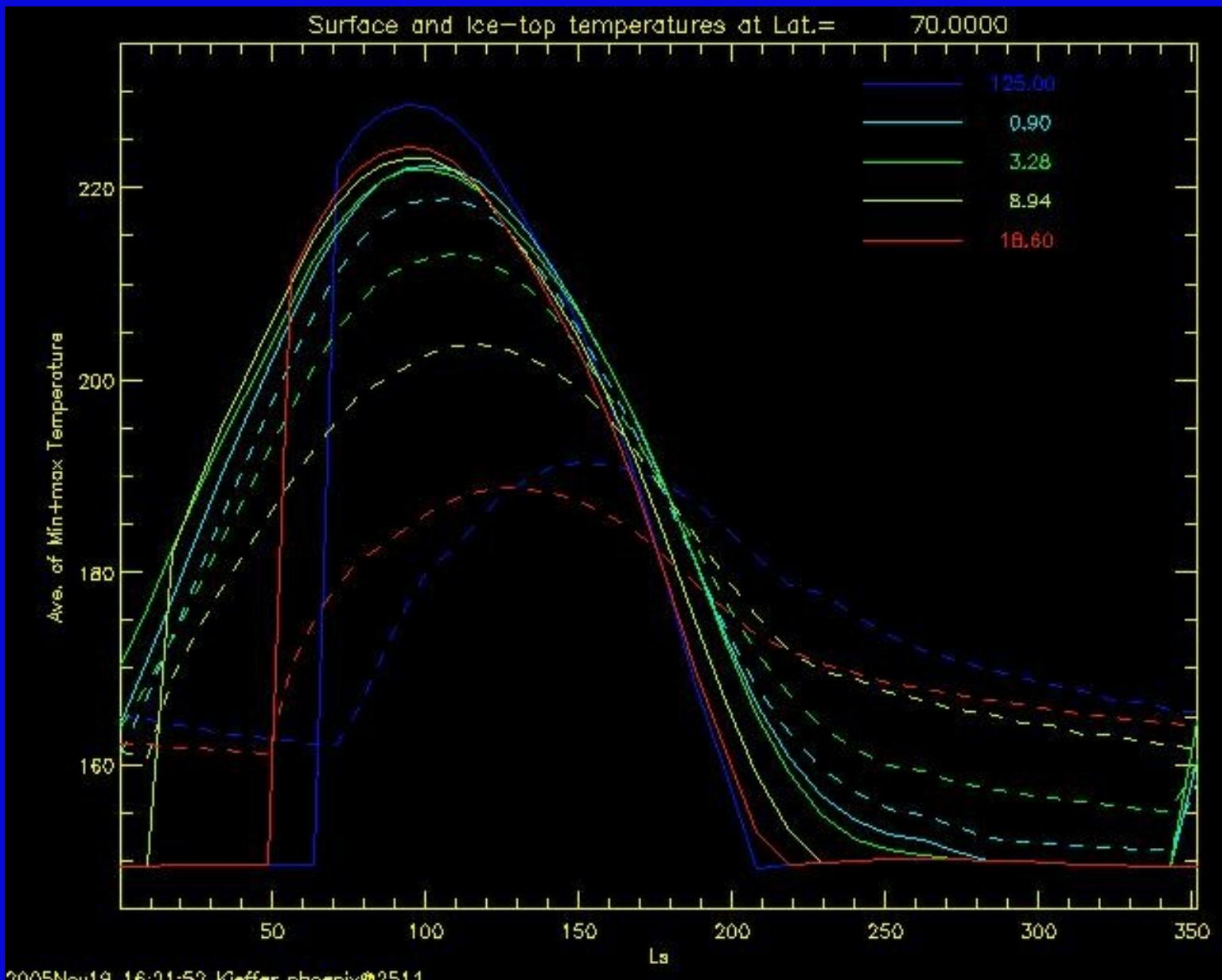
- May be a challenge to retain flexibility

Extras

Evolution of Mars' thermal instruments



KRC: Effect of depth to ice

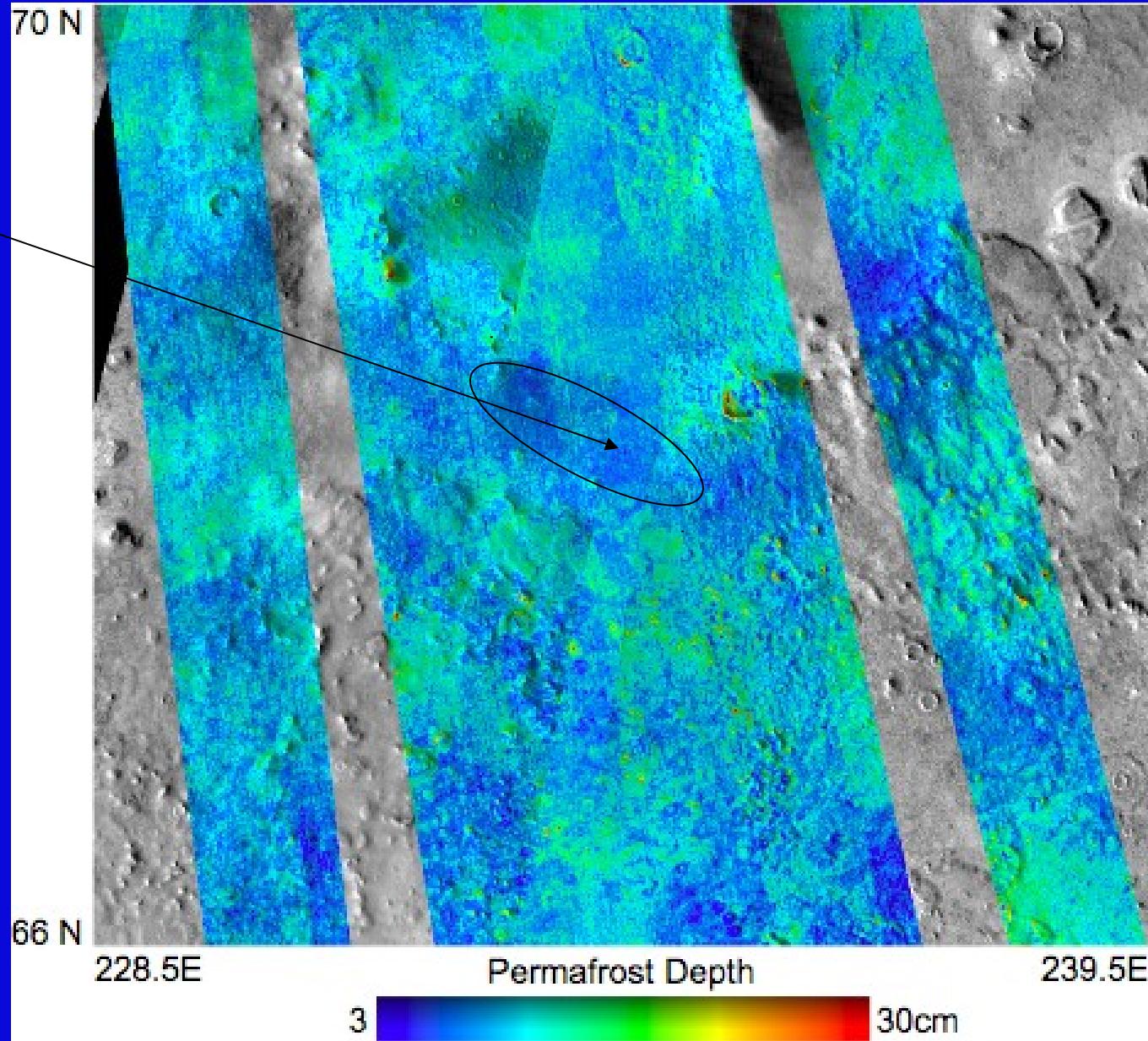


Bandfield et al. THEMIS mapping of landing region

Phoenix
Landing
site

Regional:
TES:
 $I=258$,
 $Z=4.5\text{cm}$

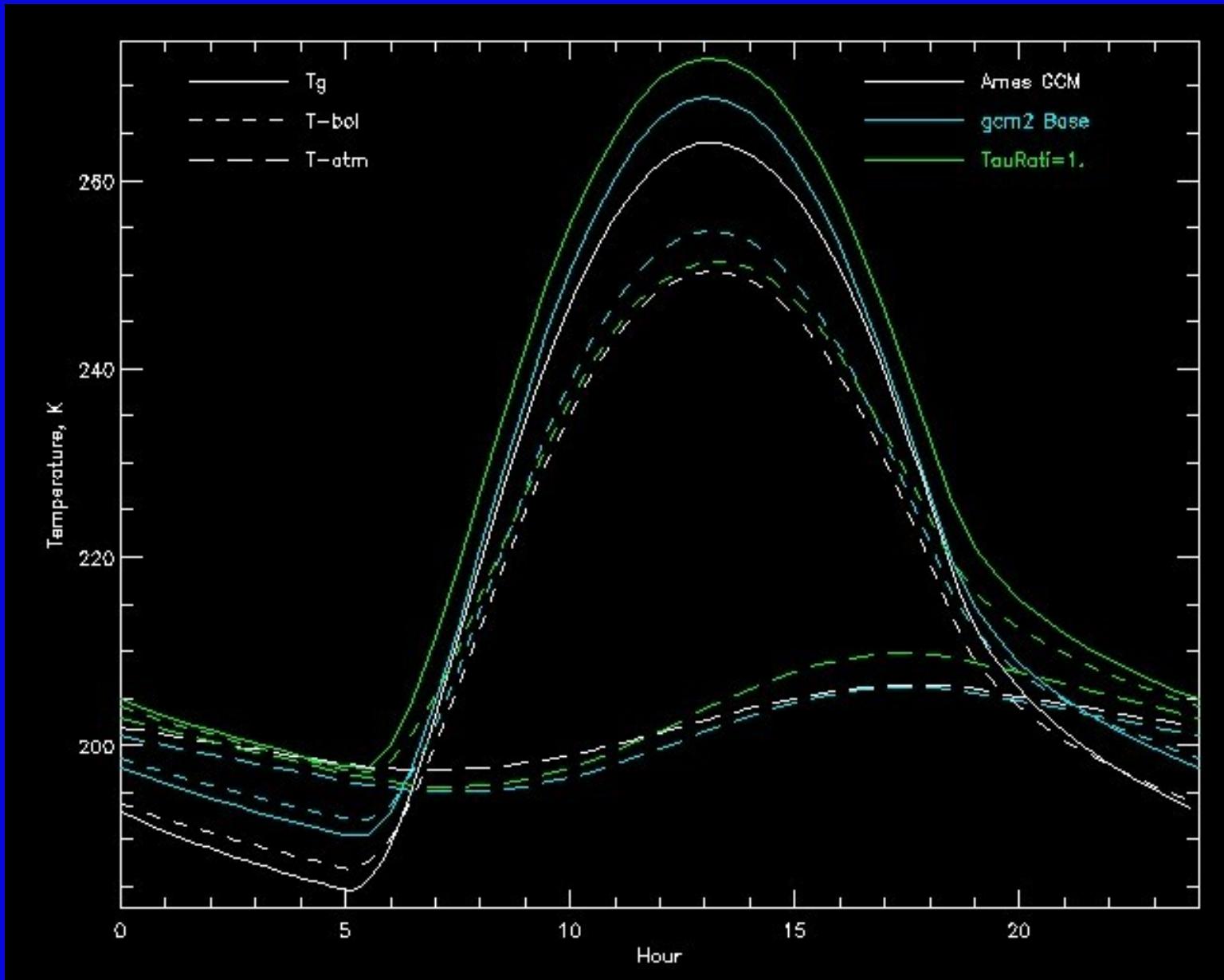
THEMIS:
 $I=283$
 $Z=6.2\text{cm}$



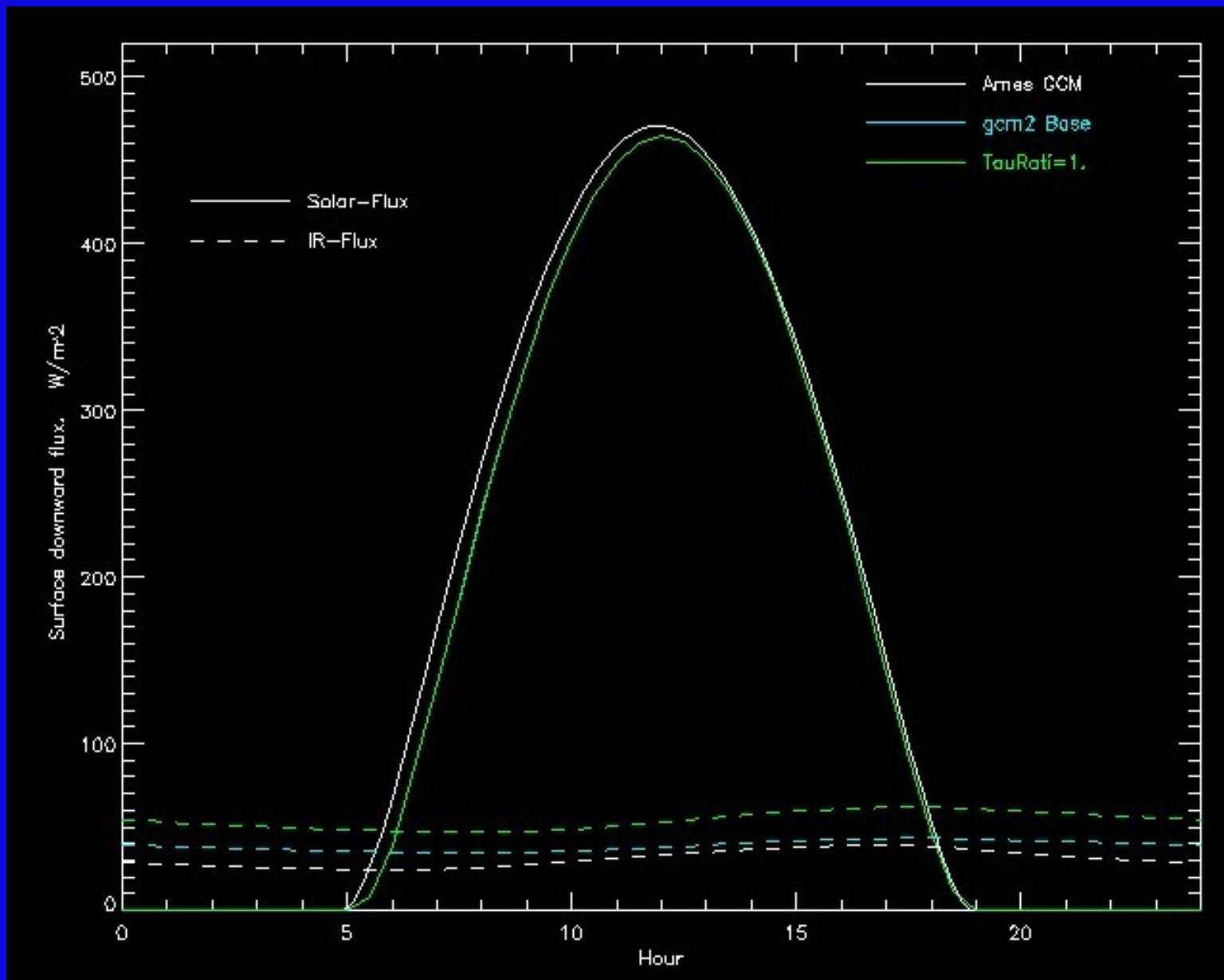
Mix of Rocks and fines is common

- Rocks dominate at night
 - Anything larger than ~10 cm
- Fine particles dominate in day
- Using day and night observations allow determination of **rock fraction** and drift particle size.
- Good predictions for landing sites
- With spectral information, need to weight radiances, not temperatures
- KRC is for one vertical situation at a time.
- Linear addition of models represent scenes of mixed rocks and fines
- Mixed scenes do not look like blackbodies; brightness temperature increases toward shorter wavelengths
- Day+night observations [or spectral analysis] allow separation of rocks from fines.
- Rock fraction is a major criterion for landing sites.

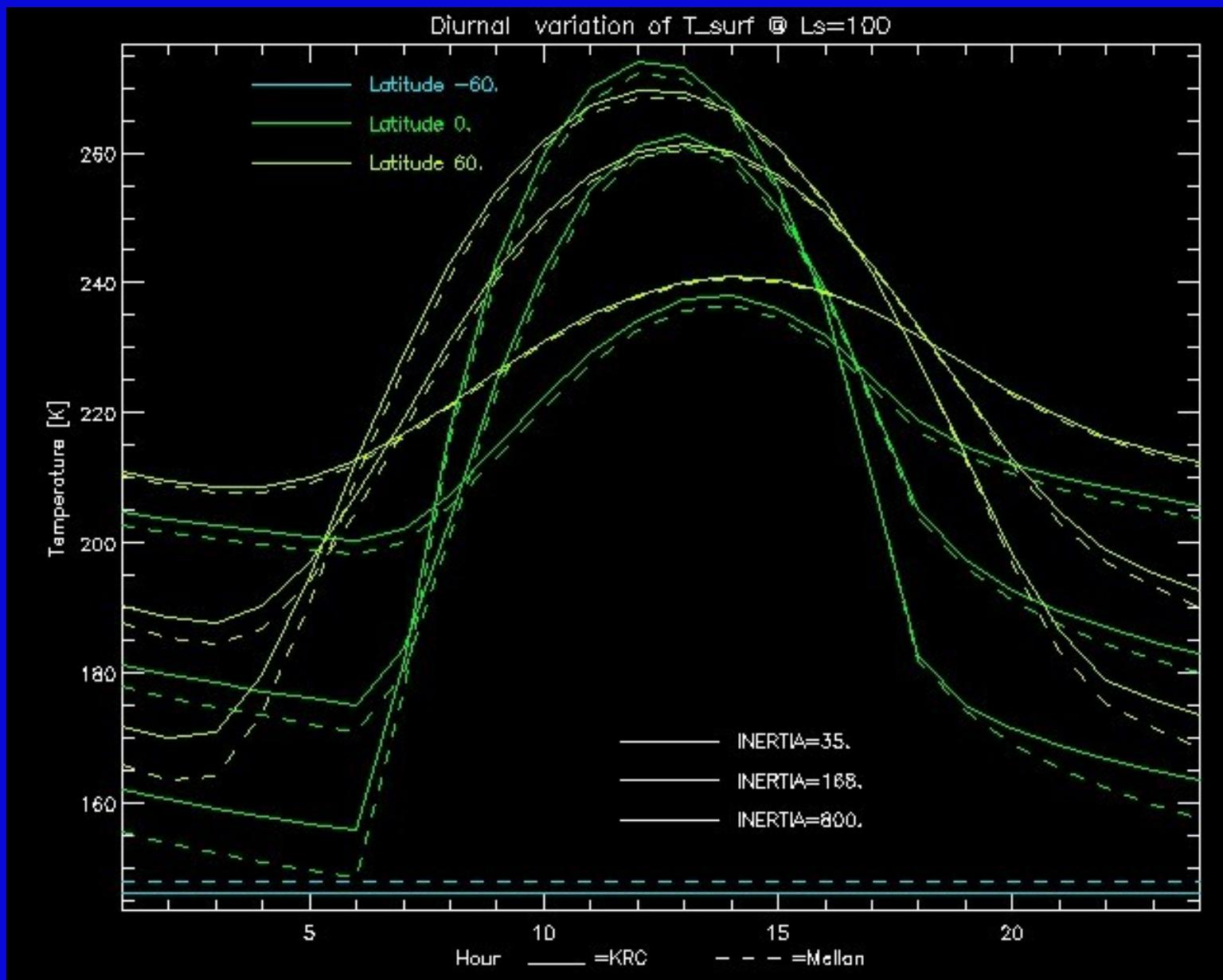
Comparison to Ames GCM model: temperatures



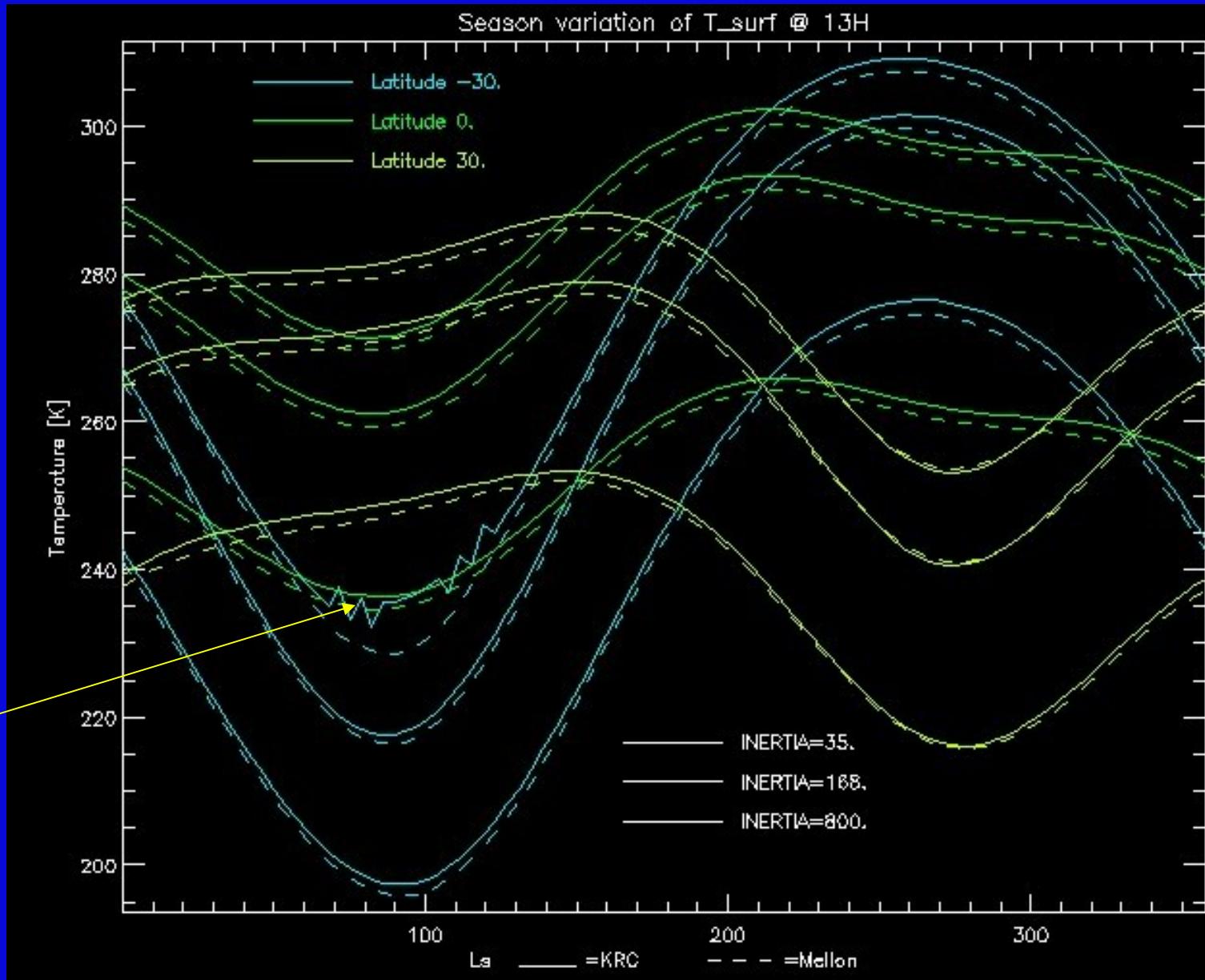
Comparison to Ames GCM model: downward flux



Comparison to Mellon Models (TES default): diurnal



Comparison to Mellon Models: seasonal



Uncertainty of Thermal Inertia

- Deal with the uncertainty of measured temperature, as well as all of the parameters that go into the model.
- For simplicity, assume the “errors” are uncorrelated
- $U_I = \text{RSS} [UT_b, UA \cdot \partial T_m / \partial A, U \tau \cdot \partial T_m / \partial \tau, U \Downarrow \cdot \partial T_m / \partial \Downarrow, \dots] / (\partial T_m / \partial I)$ at proper I

U = uncertainty A =Albedo τ =Opacity \Downarrow =slope and azimuth

All the partial derivatives are available from a thermal model!

Run models with each U parameter individually perturbed

Or, subtract pairs of models in the large sets.

You have to assess the magnitude of the input Uncertainties